



The SuperCDMS SNOLAB Experiment

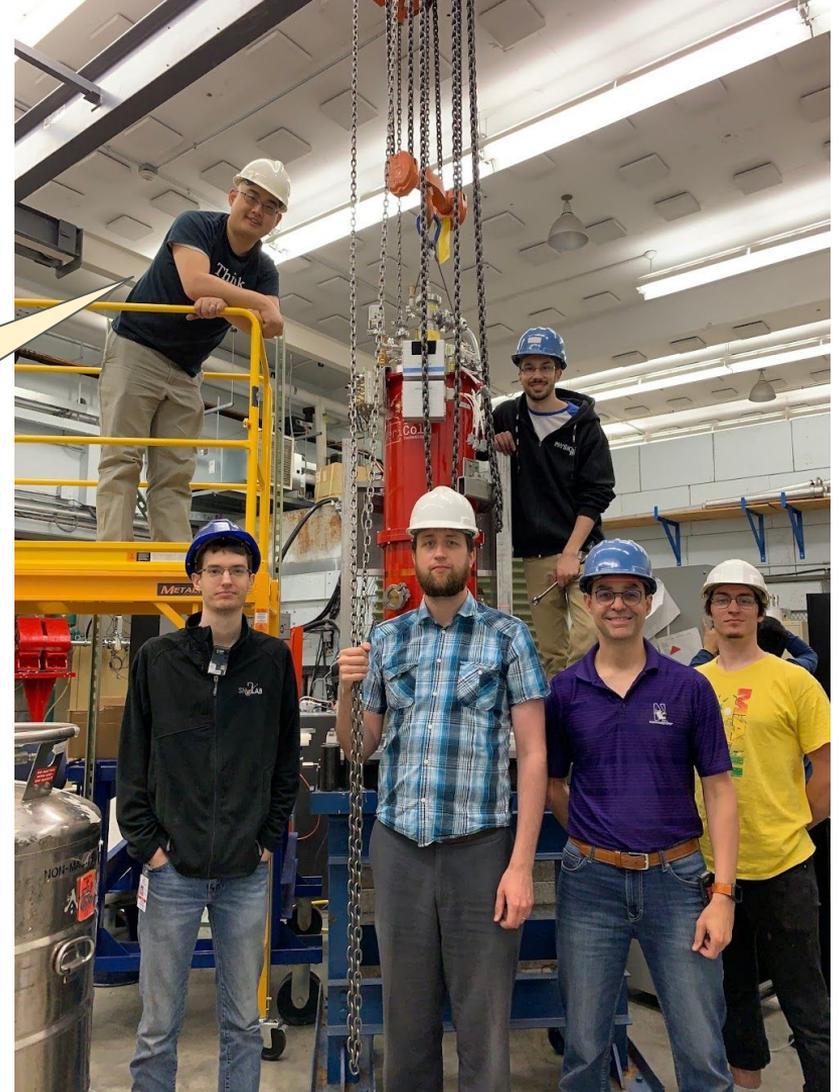
SNOLAB User's Meeting
Ziqing Hong, for the SuperCDMS Collaboration



UNIVERSITY OF
TORONTO

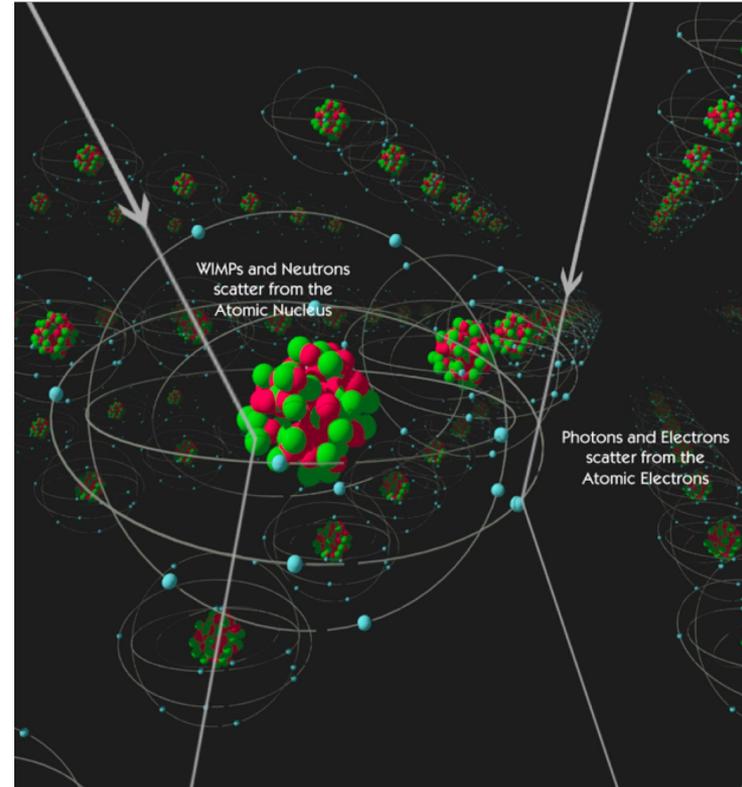


- New faculty member at University of Toronto
- Interested in Dark Matter search, Coherent Neutrino Scattering, or anything that could potentially break the Standard Model
- Specialized in cryogenic systems, cryo detector R&D and calibration
- Also love burning CPU hours with simulations...



Dark Matter Interactions & Detector Physics

- Interactions fall into two categories
 - **Nuclear recoils** - particle interacts with the nucleus
 - Traditional “WIMP” Dark matter signal
 - Neutron and neutrino backgrounds
 - **Electron recoils** - particle interacts with atomic electrons
 - Electron recoil signals and Dark Photon and Axion-like particle absorption signals
 - Most background sources
- Detector response often is different for the two categories.
 - Can be used to reject some backgrounds
- Energy scale of the interaction dictates detection approach
- Backgrounds and detection techniques drive science reach



SuperCDMS Detector Technology

Discriminating

iZIP Detector:

- Prompt phonon and ionization signals allow for discrimination between nuclear and electron recoil events

Low Threshold

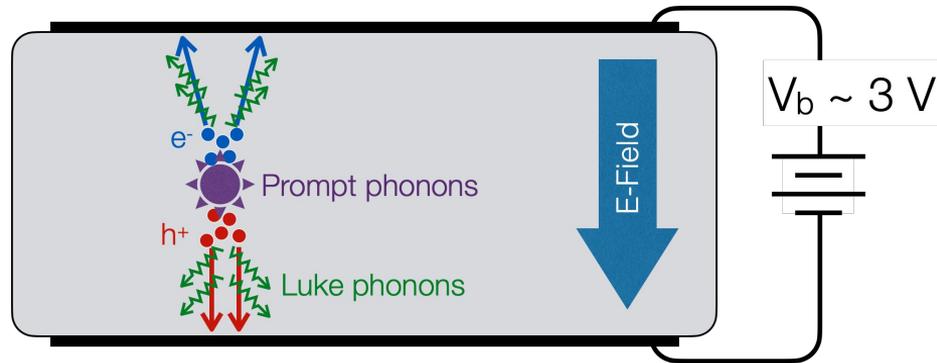
HV detector:

- Drifting electrons/holes across a potential (V_b) generates a large number of phonons (Luke phonons).
- Enables very low thresholds!
- Trade-off: No event-by-event NR/ER discrimination

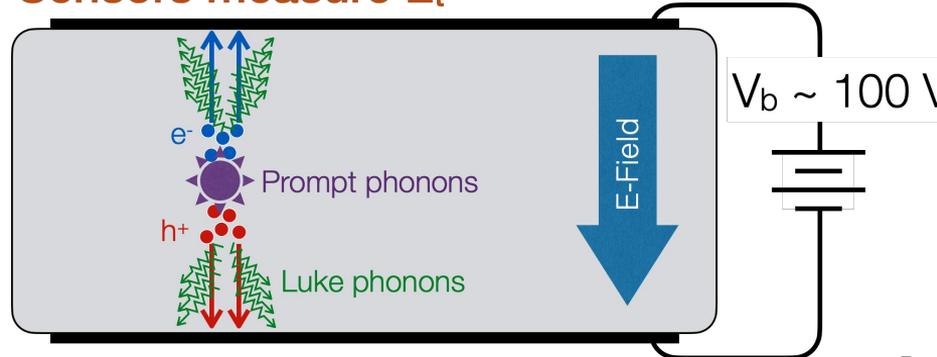
$$E_t = E_r + N_{eh} e V_b$$

E_t : total phonon energy
 E_r : primary recoil energy
 $N_{eh} e V_b$: Luke phonon energy

Sensors measure E_t , and n_{eh}

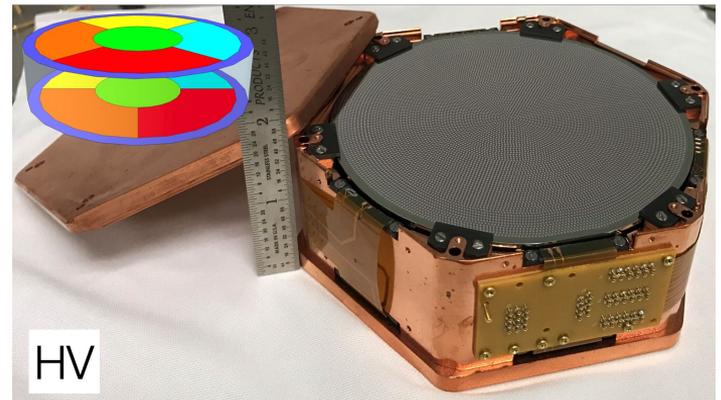
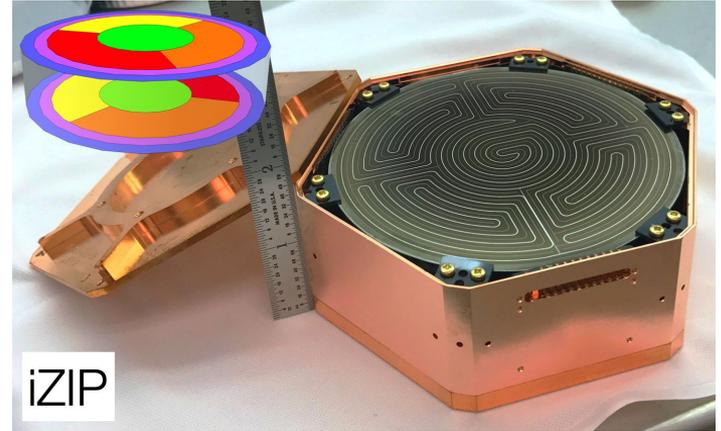


Sensors measure E_t



SuperCDMS Detectors: Posing for the Cameras

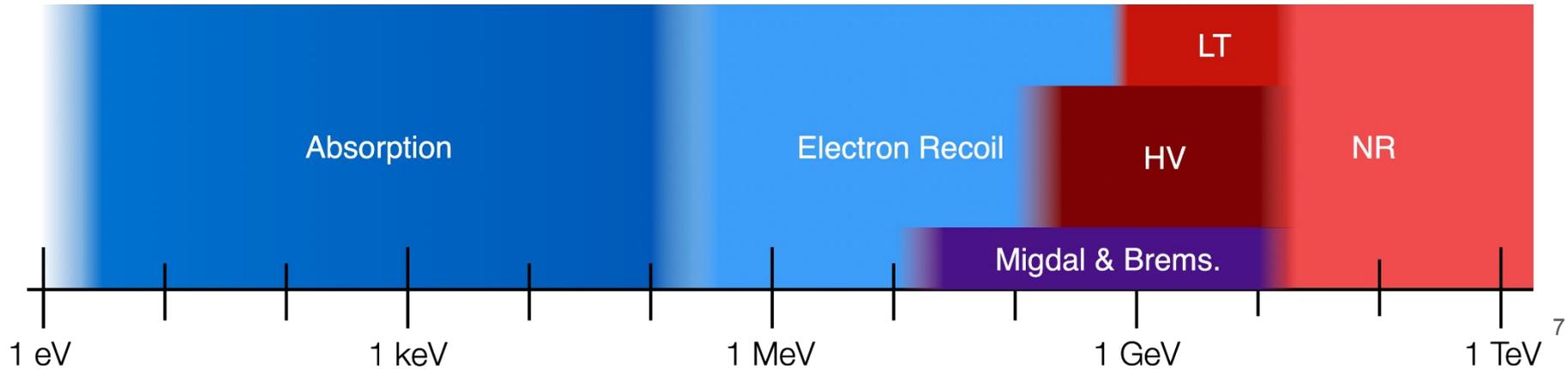
- Detectors made of high-purity Ge and Si Crystals
 - Si (0.6 kg each) provides sensitivity to lower dark matter masses
 - Ge (1.5 kg each) provides sensitivity to lower dark matter cross-sections
- Low operation temperature: $\sim 15\text{mK}$
 - Athermal phonon measurement with TESs
 - Ionization measurement (iZIP) with HEMTs
- Multiple channels per detector to identify event position
- Initial payload will consist of 4 stacks of six detectors (“towers”)
 - 2 iZIP: 10 Ge / 2 Si
 - 2 HV: 4 Ge / 8 Si



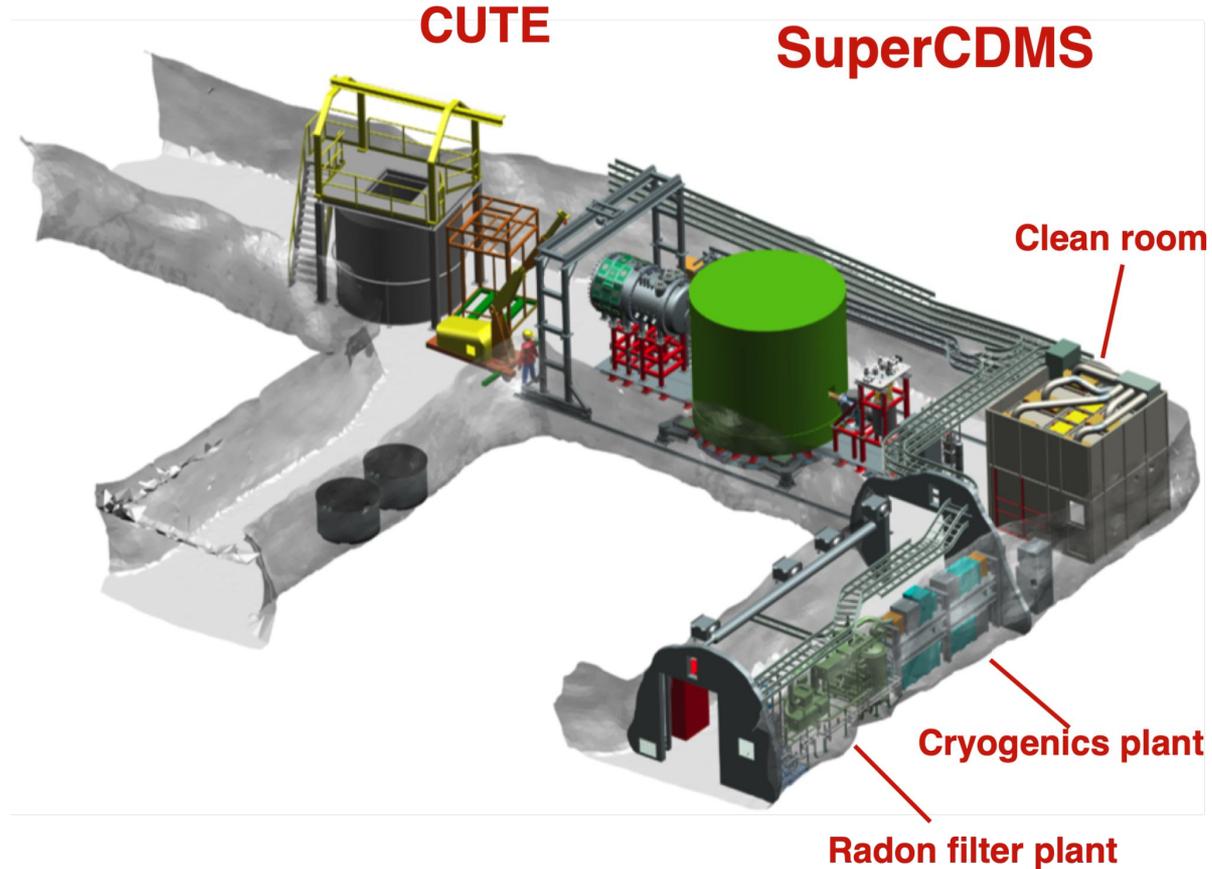
SuperCDMS Detectors & Dark Matter Mass Scales

- Dark Matter Mass Ranges

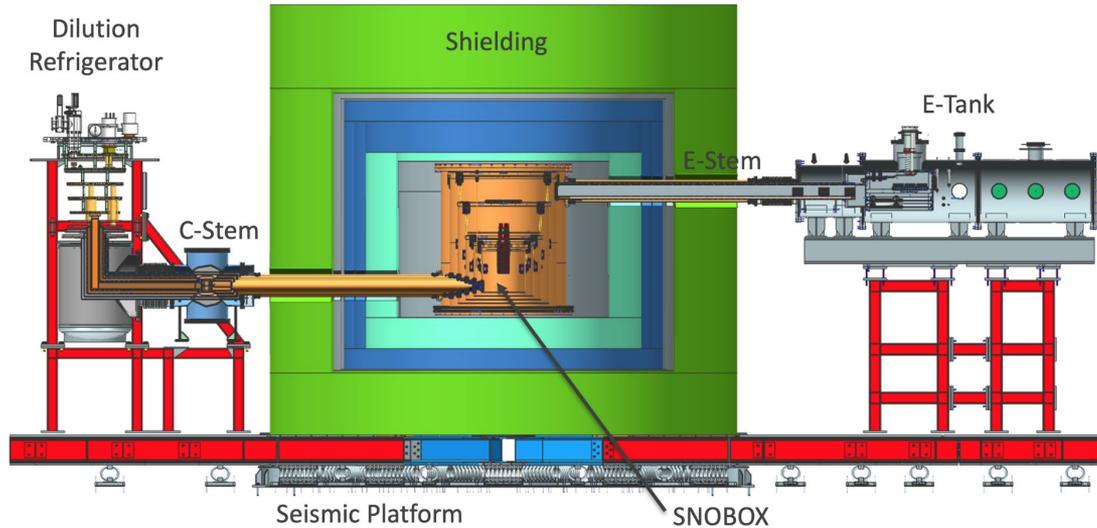
- "Traditional" Nuclear Recoil: Full discrimination, $\gtrsim 5$ GeV
- Low Threshold NR: Limited discrimination, $\gtrsim 1$ GeV
- HV Detector: HV, no discrimination, $\sim 0.3 - 10$ GeV
- Migdal & Bremsstrahlung: no discrimination, $\sim 0.01 - 10$ GeV
- Electron recoil: HV, no discrimination, ~ 0.5 MeV – 10 GeV
- Absorption (Dark Photons, ALPs): HV, no discrimination, ~ 1 eV – 500 keV ("peak search")



SuperCDMS @ SNOLAB



The SuperCDMS SNOLAB Experiment



Facility:

- 6800 m.w.e. overburden
- 15 mK base temperature
- Initial Payload: ~30 kg total
4 towers (2 iZIP, 2 HV)

Electron Recoil Backgrounds:

- External and facility: $O(0.1 / \text{keV/kg/d})$
- Det. setup: $O(0.1(\text{Ge})-1(\text{Si}) / \text{keV/kg/d})$
- Total: $O(0.1-1 / \text{keV/kg/d})$

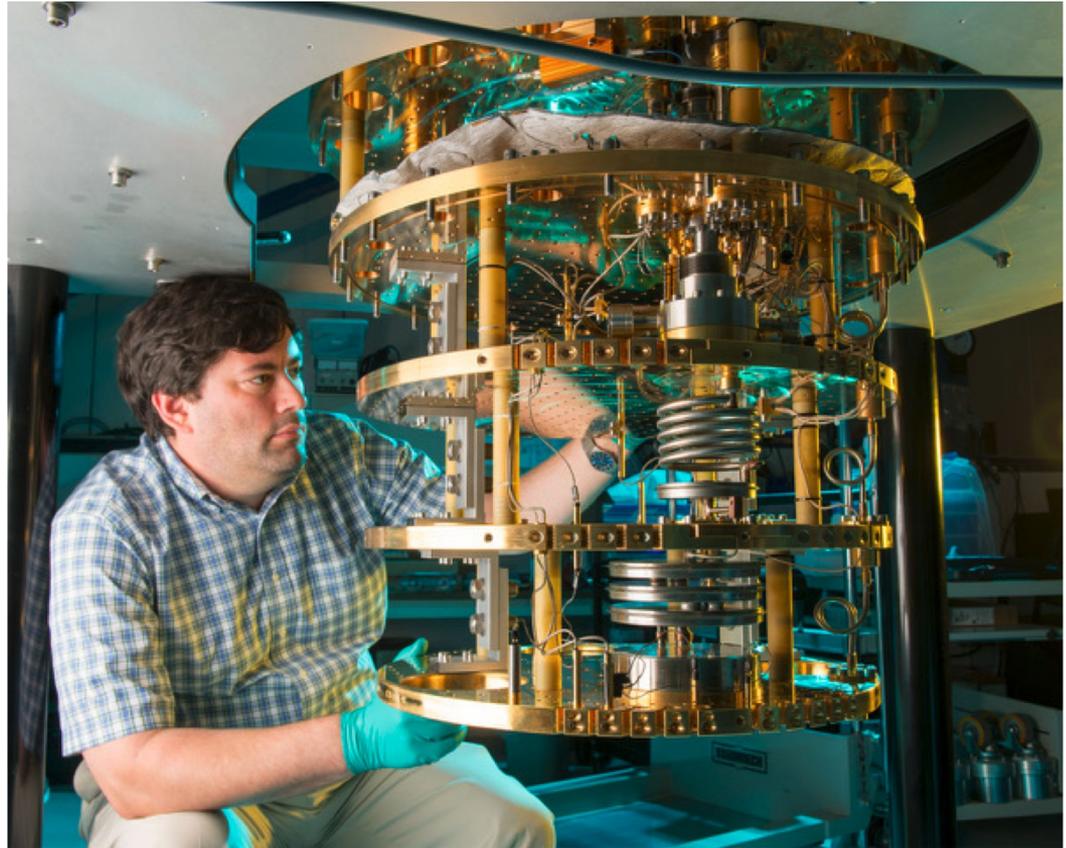
Facility designed to be dominated by solar neutrinos in NR background

Vibration isolation:

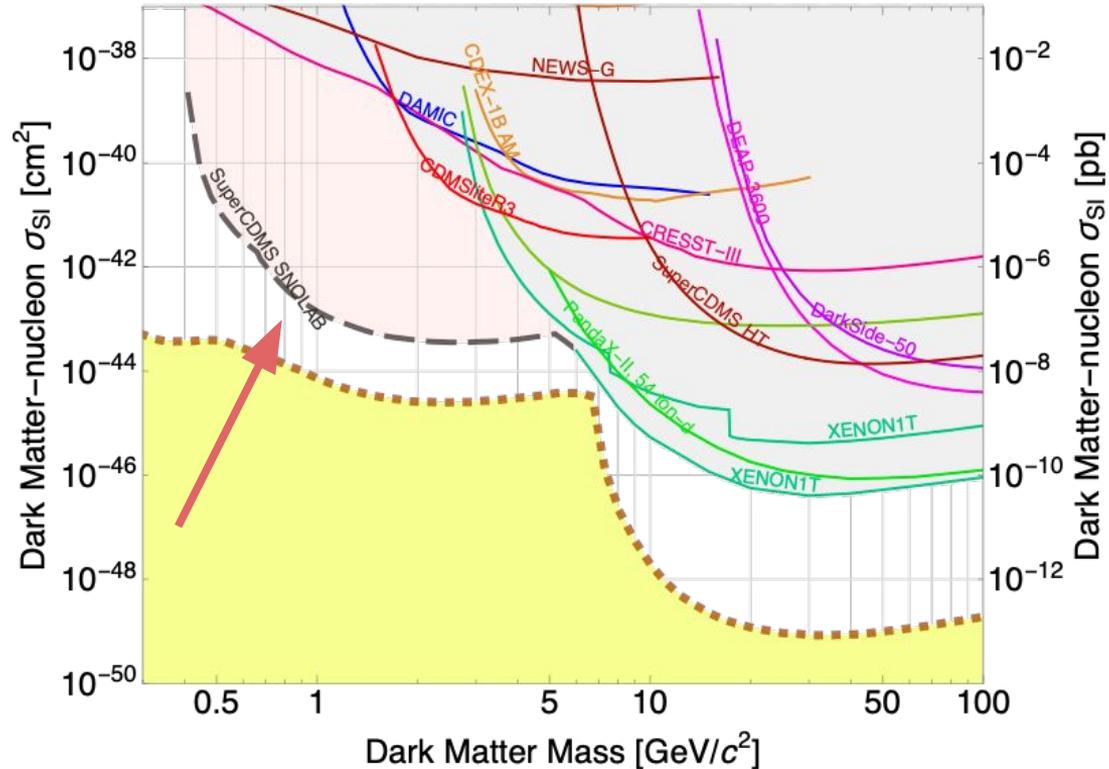
- Seismic: spring loaded platform
- Cryo coolers: soft couplings
 - Braids, bellows
- Copper cans: hanging on Kevlar ropes

The SuperCDMS Dilution Refrigerator

- Big machine!
- Will show up at SNOLAB at the beginning of 2022!
- The rest of the cryostat will follow and get integrated
- Full cryogenic system commissioning early 2023
- Expect cold detectors fall 2023

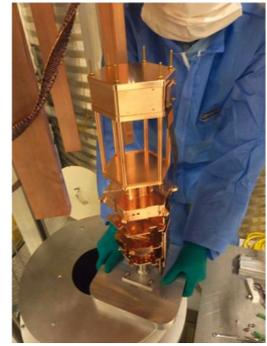


SuperCDMS SNOLAB Projected Sensitivity

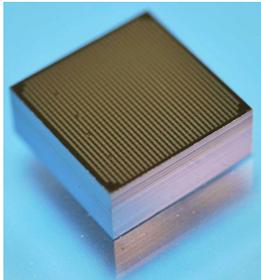


R&D and Detector Testing in Parallel

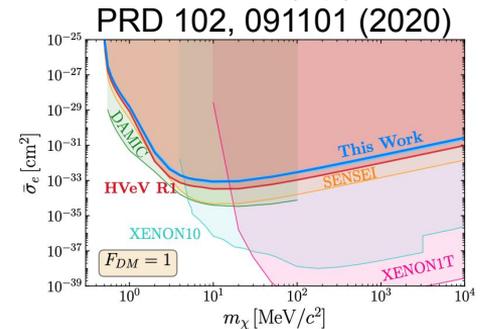
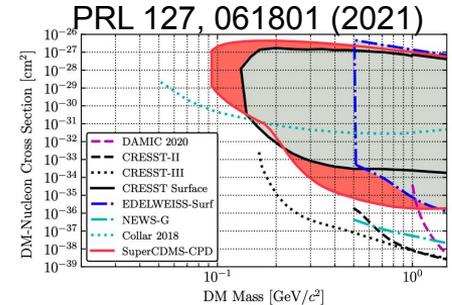
- Testing prototype detectors at CUTE
 - Please refer to the CUTE status talk by Silvia yesterday
- R&D of smaller devices achieving new sensitivities to low mass DM models



10 gram silicon
3.9 eV resolution
0 V operation

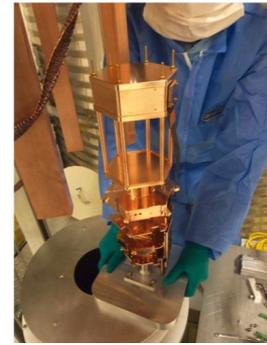


1 gram silicon
2.6 eV resolution
100 V operation
→ 0.03 charge resolution



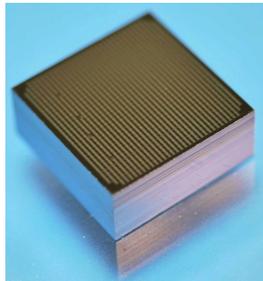
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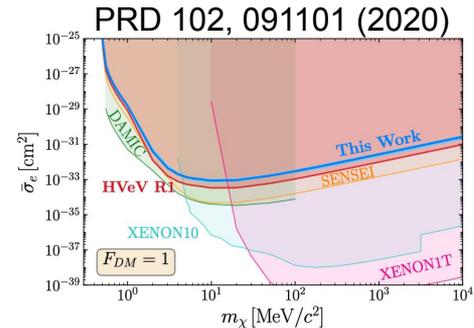
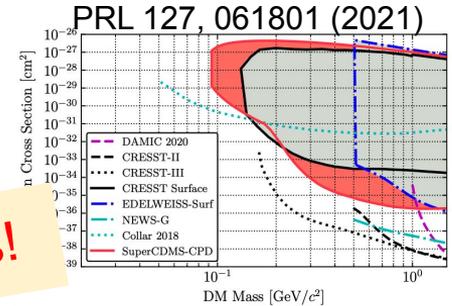


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Candidates for future upgrades!

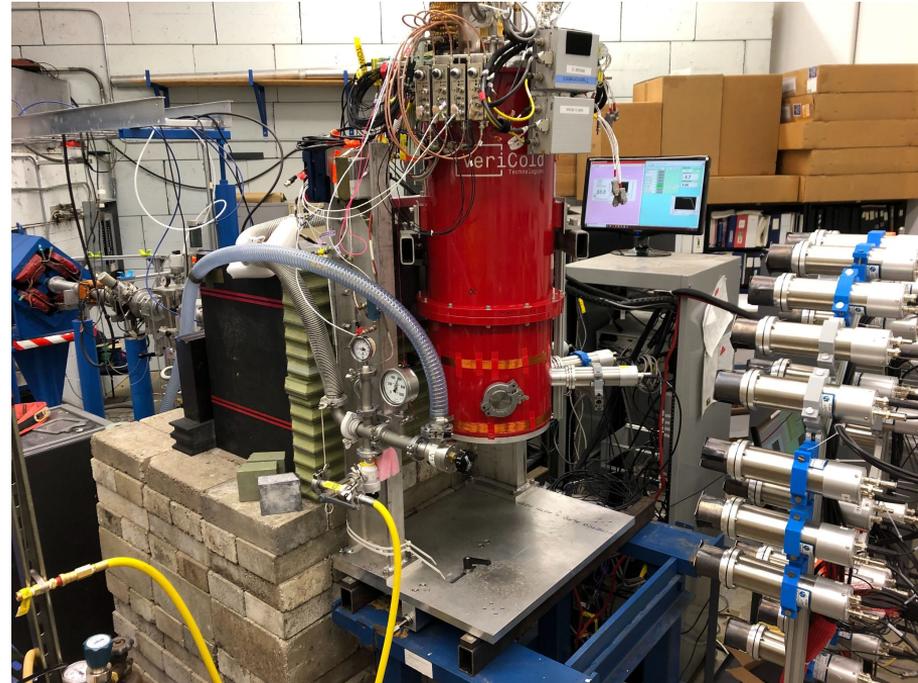


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Ionization yield measurement for Si and Ge

- Measuring ionization yield of Si and Ge
 - At energy below 1 keV
 - To be ready for data interpretation as soon as the SNOLAB cryostat turns on
- Classical scattering experiment
 - A monoenergetic neutron beam
 - A Si/Ge detector in a mobile cryostat
 - An array of secondary neutron detectors to tag the recoil angles
- First Si measurement done in 2019
 - Result coming out this fall
- Ge measurement planned for 2022
 - A Canadian-led effort now



Conclusions

- SuperCDMS detectors aiming to reach “neutrino floor” in 1-10 GeV NR mass range
- Technology being adapted in smaller detectors to search for light dark matter, down to
 - $\mathcal{O}(10)$ MeV via inelastic Nuclear recoil channels (Migdal, Bremsstrahlung)
 - $\mathcal{O}(1)$ MeV via Electron recoil channels and
 - $\mathcal{O}(1)$ eV via Dark Photon and Axion-like Particle (ALP) Absorption channels
- SuperCDMS designed a powerful complex cryogenic system that is being installed at SNOLAB
 - SuperCDMS Detector installation – next spring/summer
 - Full system commissioning – beginning of 2023
 - Cold detector – fall 2023
- SuperCDMS is particularly competitive at low masses, including electronic interactions.
- R&D on small/more sensitive devices as well as ionization yield measurements are progressing
- Stay tuned! Experiments are producing results at a fast pace, more sensitive experiments are soon to come online.