# The New SNOLAB Underground Low Background Laboratory (LBL)

Richard Ford and Ian Lawson (SNOLAB) For the SEF Meeting (Toronto) 9<sup>th</sup> March 2017

# **Brief Outline**

- Current status
- Current detectors and capabilities
- Future capabilities
- Future into the future plans
- AARM and Community Assay Database



### **SNOLAB Underground Facilities**



# South Drift Today







### South Drift Plan for LBL



## General Purpose Underground Shielding Tank (GUST??)



- Free standing tank with integral deck
- 8 m diameter, 8 m high
- Walls constructed from bolted, corrugated cylindrical segments of galvanized carbon steel
- Polyurea lined tank ~ 1 cm thick coats walls, floor, and transition to ss lid
- O-ring sealed nozzles and penetrations
- Overfilled tank (~ 4 inches above tank lid bottom)



### **Current Capabilities**

- PGT Ge Detector
  - 210 cm<sup>3</sup>, 83mm endcap, 1.8 KeV FWHM, shield 2" Cu + 8" Pb
- Canberra Well Ge Detector
  - 300 cm<sup>3</sup> with 3ml well sample volume
- Canberra Large Ge Detector
  - Coaxial 400 cm<sup>3</sup>, shield 3" Cu + 4" LB-PB + 6" Pb
- Gopher HPGe Detector (SuperCDMS)
  - 2.0 Kg P-type coaxial, 2" OFHC Cu + 2" LB-Pb + 8" Pb
- XIA Alpha Detector
  - UltraLo-1800, < 0.0001 alpha/cm2/hr
- VDA Ge Detector (EXO)
  - Coaxial from U. Bern, 400 cm3, shield 6" Cu + 6" Pb
- XRF detector
  - From SNO, developed for dust sample counting (Fe, Ca, Zn)
- Tennelec Alpha Counters
  - With front glove box
- Beta-Alpha PSD Coincidence Counters (currently on surface)
  - From SNO developed for Ra counting in liquid scintillator
- Electrostatic Counters (ESCs) (currently on surface)
  - With low pressure N2 recirculation for Rn and Ra counting.
- Radon Emanation and Extraction Board
  - Small Marinelli shaped chamber to be rebuild with large chamber

# **SNOLAB PGT HPGe Counter**

(The workhorse detector at SNOLAB)



## **Unshielded and Shielded Spectra**

(PGT Coax Detector)

Energy (keV)



## **PGT HPGe Typical Detector Sensitivity**

(for a standard 1L or 1 kg sample counted for one week)

Isotope	Sensitivity for Standard Size Samples
238 U	0.12 mBq
235 U 232	0.17 mBq
Th	0.11 mBq
<sup>40</sup> K	1.50 mBq
60 Co 137	0.05 mBq
Cs	0.14 mBq
Mn	0.05 mBq

## **Canberra Well Detector at SNOLAB**



## **Canberra Well Detector at SNOLAB**



Detector Volume: 300 cm<sup>3</sup>

Sample Well

### Typical Sample Containers

### Volume is 3 ml

Volume is 8 ml



## **Unshielded and Shielded Spectra**

(Canberra Well Detector)

Energy (keV)



## **Canberra Well Detector Sensitivity**

lsotope	Sensitivity for Standard Size Samples
<sup>238</sup> 226	0.04 mBq
U (↓ Ra)	0.03 mBq
	0.12 mBq
Th	0.23 mBq
235 U	0.01 mBq
Pb	0.08 mBq

## **Gopher HPGe**

2.0kg of Ge. P-type coaxial

- Canberra Ge detector with thin 1.6 mm Al window, allows sensitivity to 210Pb
- Detector is 400 cm3 and is shielded by
  - 2" inner OFHC copper
  - Surrounded by 2" low-activty Pb
  - Surrounded by > 10" normal Pb
  - Plus some outer polyethelene
- The shielding box is purged with 1.8
- Derdicateidrogsopeoidentiseas.
- Sensitivity of ~ 1 mBq/kg for 3 week run
- Sample changes by SNOLAB scientists and staff
- Queue and Analysis by SuperCDMS
- Current status: Background run in progress



## Vue des Alpes HPGe Detector

The VdA HPGE detector has been reconditioned by baking and vacuum pumping for several months.

Calibration runs have been done to verify peak resolution and now a longterm background run is in progress.







## Canberra Coaxial HPGe Detector

Canberra 400 cm<sup>3</sup> p-type coaxial HPGE detector acquired in 2011.

Refurbished into an ultra-low detector in 2013 with verified low-background materials.

Shielding is partially assembled.

Machining of the lead bricks is underway for the cold finger and nitrogen purge lines.

Construction of the shielding plug is underway.





# **Electrostatic Counting System (ESCs)**



9 counters located at SNOLAB,1 on loan to LBL (EXO),1 on loan to U of A (DEAP).

Originally built for SNO, now used primarily by EXO. However, these counters are owned by SNOLAB so samples can be measured for other experiments.

Measures <sup>222</sup>Rn, <sup>224</sup>Ra and <sup>226</sup>Ra levels. The technique involves recirculation of low pressure gas from sample volume to the ESC.

Sensitivity Levels are:

<sup>222</sup>Rn: 10<sup>-14</sup> gU/g <sup>224</sup>Ra: 10<sup>-15</sup> gTh/g

<sup>226</sup>Ra: 10<sup>-16</sup> gU/g

Work is ongoing to improve sensitivity even further.

# **Alpha Beta Counting System**



Transparent liquid scintillator vials optically coupled to 2" PMTs.

The technique is combination of pulse shape discrimination and coincidence counting for identifying BiPo events.

Sensitivity for  $^{238}$ U and  $^{232}$ Th is ~1 mBq assuming that the chains are in equilibrium.



# **XIA Alpha Counter**



# **Radon Emanation Chamber**



## Future Plans for SNOLAB LBL

- Extended use for the DUST tank as general purpose shielding tank with veto. For example, may install acrylic scintillator tank and PMT array for counter.
- Low radon air supply.
- Low radon air room for assembly and sample preparation.
- New Radon emanation capabilities, eg. large emanation tanks and system to extract from experiment process systems.
- New detectors as needed and motivated:
  - Alpha counters
  - XIA high sensitivity large area surface alpha counter
  - Beta cage counter
- New chemical analytical capability, eg. HP-ICPMS.
- Improved liquid particle counting.

# Low Radon Air Supply

- To supply to assembly rooms, detectors or gloveboxes
- General methods are low-pressure adsorption, or pressure/vacuum swing adsorption, or a hybrid or these.
- Flow rates over 50 m<sup>3</sup>/hr and radon levels below 100 mBq/m<sup>3</sup> have been demonstrated.
- Maybe able to use Vale mine compressed air system as input air, which is surface air (3 – 9 Bq/m<sup>3</sup>), versus UG air ~130 Bq/m<sup>3</sup>.



## SNOLAB is a member of AARM:



## Communication and Collaboration on common issues related to Assay and Acquisition of Radiopure Materials

Content and management of the AARM website to be moved to LRT website www.lowradioactivity.org

## **The AARM Integration Website**

### Home Page

### Assay and Acquisition of Radiopure Materials

a cooperative approach to low radiation techniques and sensitive background measurement



#### Home

#### Community

Facilities and Scheduling:

- Detectors info
- Assay request
- Storage request

Community Assay Database

Background Studies

Laboratory Materials

Clean-handling Procedures

Resources:

Publications

AARM presentations

LRT presentations

Go to Lowrad Wiki

Interdisciplinary Research Portais The purpose of this website is to share knowledge and resources across the community of those involved in low-radioactivity experimental techniques. The tasks covered include

- Coordinating users and providers of sensitive assay facilities
- Linking to information on availability of radiopure materials
- Sharing knowledge concerning backgrounds in low-background experiments

#### NEWS:

The next AARM meeting is May 20-21, 2015. It will be held right after the Conference on Science at SURF hosted at the South Dakota School of Mines & Technology.

Summary agenda on the AARM wiki and copied below. For more detail see the full agenda.

#### May 20 (Wednesday)

12:30-1:45 Discussion over box lunch:

- AARM (website and other common tools)
- G2 collaboration infrastructure priorities
- 0vββ background infrastructure



### **The AARM Integration Website**

AAF

Assa	Facilities and Scheduling: Assay or Storage Request What the user submits
	Check out the available assay from Detector Info
Home	Fill out request form – goes to the contact listed
Community Facilities ar	Storage request also available
Detectors info Assay request	Click here to request login information For general information on proper handling of samples refer to the Clean-handling procedures.
Storage request	For information on the results of previous assays please see the Community Assay Database
Database Background Studi	es *Login: *Password:
Laboratory Materi Clean-handling Procedures	als     *Requester Name:       *email address:
Resources: Publications	*Experiment:
AARM presentat	ions     Number of samples:       ns     Please describe sample size, shape, mass etc. here.
Go to Lowrad Wik Interdisciplinary	i Sample dimensions:
Research Portals	Sample composition:
	Sample description:

Suggestions? Send them to fritts@physics.umn.edu and astrong@umn.edu

### The AARM Integration Website

### **Resources:**

Easy reference for relevant

### papers

### Links to LRT and AARM

### presentations

### Add your own naner

#### Home

Community

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Go to Lov

Interdisciplinary Research Portals

luminum as a source of background in low background e	x
. Majorovits, I. Abt, M. Laubenstein, O. Volynets.	

[arXiv:1105.3591 [physics.ins-det]].

Cosmogenic Production as a Background in Searching for R D.-M. Mei, Z.-B. Yin, S.R. Elliott. [arXiv:0903.2273 [nucl-ex]].

Cosmogenic activation in germanium and copper for rare e S. Cebrian, H. Gomez, G. Luzon, J. Morales, A. Tomas, J.A.

Cosmogenic activation of a natural tellurium target. V. Lozza, J. Petzoldt. [arXiv:1411.5947 [physics.ins-det]].

Measurement of the specific activity of ar-39 in natural argo WARP Collaboration (P. Benetti et al.). [astro-ph/0603131].

#### Geant / FLUKA

Cosmogenic Backgrounds in Borexino at 3800 m water-equivalent depth. Borexino Collaboration (G. Bellini et al.). [arXiv:1304.7381 [physics.ins-det]].

Monte Carlo Simulation of HERD Calorimeter. M. Xu, G.M. Chen, Y.W. Dong, J.G. Lu, Z. Quan, L. Wang, Z.G. Wang, B.B. Wu et al.. [arXiv:1407.4530 [astro-ph.IM]].

Muon-induced neutron production and detection with GEANT4 and FLUKA. H.M. Araujo, V.A. Kudryavtsev, N.J.C. Spooner, T.J. Sumner. [hep-ex/0411026].

Suggestions? Send them to fritts@physics.umn.edu and astrong@umn.edu

#### Presentations from Low Radiation Techniques (LRT) we

The Low Radioactivity Techniques (LRT) workshop series examines topics in low radioactivity techniques, a fundamental aspect of rare event searches. Topics include dark matter, solar i decay, and long half-life phenomena. This conference's wide scope includes all aspects of the background detectors and techniques.

The goal of this workshop series is to bring together experts in this field for presentations ar broadly the issues of low radioactivity techniques. The intention is to foster and continue the resource sharing required for new generations of detectors to be developed at underground

The LRT workshop series started in 2001 at Laurentian University in Sudbury. Presentations 2004 are available at the links below:

#### LRT 2015: Presentations LRT 2013: Presentations LRT 2010: Presentations LRT 2006: Presentations LRT 2004: Presentations

#### Presentations from AARM meetings

The original purpose of the Acquisition and Assay of Radiop Facility for AARM (or FAARM) at DUSEL. This effort was fund Tools for Underground Science", was funded beginning in 20 background community, as represented by this website.

The AARM group has held meetings since 2010 to define go meetings are available at the links below:

March 19-22, 2014 at Fermilab

March 4, 2013 at SLAC National Accelerator Laboratory

June 22-23, 2012 at University of Minnesota

November 11-12, 2011 at University of Minnesota

February 25-26, 2011 at University of Minnesota

November 12-13, 2010 at Sanford Underground Research

March 19-20, 2010 at UC Berkeley

## http://www.radiopurity.org/

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s			Total r	esult: 68				
als	Grouping	Name		Isotope	Amount	Isotope	Amount	
	► BOREXINO (200	2) Stainless steel for stora	ne vessel	Th-232	9 3e-10 a/a	11-238	1 4e-9 a/a	
	BOREXINO (200	<ol> <li>Stainless steel TK3B for</li> </ol>	storade vessel	Th-232	1 2e-9 a/a	11-238	4e-10 g/g	
	BOREXINO (200	<ol> <li>Stainless steel foil for R</li> </ol>	emanation test	Th-232	5e-11 a/a	11-238	5e-11 a/a	
S	BOREXINO (200	2) Stainless steel AISI304	I for SSS	Th-232	2 8e-9 a/a	11-238	3 7e-10 a/a	
	► BOREXINO (200	<ol> <li>Stainless steel, piping</li> </ol>		Th-232	2.5e-9 a/a	U-238	1.1e-9 a/a	
	► BOREXINO (200	<ol> <li>Stainless steel for fland</li> </ol>	es	Th-232	1.6e-9 g/a	U-238	5e-10 g/g	
	EDELWEISS (20	11) Mild steel. support		Th-232	0.01 ppb	U-238	0.01 ppb	
	► EXO (2008)	Polvimide substrate. Es	panex flat cable. Nipp	Th	450 ppt	U	900 ppt	
	► EXO (2008)	Copper coating. Espane	x flat cable, Nippon St	Th	69 ppt	U	100 ppt	
	► EXO (2008)	Copper coating, Espane	x flat cable, Nippon St	Th	135 ppt	U	67 ppt	
	► EXO (2008)	Polyimide substrate. Es	panex flat cable, Nipp	Th	1600 ppt	U	1500 ppt	
	EXO (2008)	Stainless steel, Helicoil		Th	320 ppt	U	193 ppt	
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Suggestions? Send them to fritts@physics.umn.edu and astrong@umn.edu and prisca@physics.umn.edu



Community Material Assay Database

AARM recognized 10 years ago the need for an Assay Database with an easy-to-use interface to enter new data and to search for radiopure materials, vendors, etc.
 The Majorana database (J. Loach) was chosen as a template

The more globally used, the more useful for everyone.

• Began by entering legacy data:

~300 assays; data from ILIAS, EXO and XENON100

- Now have ~ 1000 assays, primarily historical.
- Database is being used by SuperCDMS and DarkSide, LZ evaluating it
- In the process of adding features to the software for easier distributed use.



### Community Material Assay Database

	(pf? p	olyimide kapton -ptfe -raw) A	ND grouping:	EXO	<b>م</b> >		
		Total re	sult: 10				
Grouping	Name		Isotope	Amount	Isotope	Amount	
• EXO (2008)	Polyimide substrate, E	spanex flat cable, Nippon	Th	450 ppt	U	900 ppt	 1
• EXO (2008)	PFA, Saint Gobain sup	plied DuPont 450-HPB	Th	65 ppt	U	75 ppt	
• EXO (2008)	PFA, Saint Gobain Dui	Pont 440-HP	Th	13.3 ppt	U	3 ppt	 ,
• EXO (2008)	Polyimide substrate, E	spanex flat cable, Nippon	Th	1600 ppt	U	1500 ppt	
• EXO (2008)	Polyimide tape, Stanfo	ord stock room	Th	5400 ppt	U	5800 ppt	
• EXO (2008)	Copper coating, Espar	nex flat cable, Nippon Ste	Th	3 ppt	U	19 ppt	 •
• EXO (2008)	Sheldal superinsulation	n, DuPont Kapton alumini	Th	1540 ppt	U	2500 ppt	
• EXO (2008)	Polyimide substrate, E	spanex flat cable, Nippon	Th	50 ppt	U	450 ppt	 ,
• EXO (2008)	Polyimide substrate, E	spanex flat cable, Nippon	Th	317 ppt	U	3880 ppt	 ;
							-



### Community Material Assay Database

		(pf7	polyimide	e kapton	-ptfe	-raw) A	ND grouping	:EXO		<i>ه</i> >				
						Total res	ult: 10							
Grouping	Name						Isotope	Amou	Int	Isotope	Amount			
• EXO (2008)	Polyimide substrate, Espanex flat cable, Nippon					Th	450 ppt		U	900 ppt				
• EXO (2008)	PFA, Sa	aint Gobain si	upplied Du	Pont 450	D-HPB		Th	65 p	pt	U	75 ppt	[	p /	
	Sample	Description	Saint ( supplie	Gobain su er's defau	ipplie	d DuPon cedures	t 450-HPB P	PFA. Materia	al finish	ed using				
		ID	Table 3	3. #47										
	Measurement	Results	к		740	(77)	ppb							
						10.00								
			Th		65	(6.5)	ppt							
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# End

## **Reasons for Low Background Counting**

- Rare event detector experiments (eg. dark matter searches and neutrino detector) require low radioactivity background. This requires that fabrication materials have low concentrations of radioactivity chain elements (<sup>232</sup>Th, <sup>238</sup>U, and <sup>40</sup>K), often well below ppt levels. These backgrounds can be due to manufacturing cleanliness (inclusion of mineral dust), chemical contaminations (eg. potassium), or intrinsic to materials (eg. steel).
- These levels are below that generally accessible by chemical analytical techniques, and assay method is often by radiation counting.
- Another application is counting of detector target fluids, either as part of a QC requirement, diagnostics, or purification testing.
- Backgrounds of interest can be alphas, betas, gammas or neutrons, depending on the experiment and the component in question. The radiation could be direct, or via leaching or emanation. Thus several detector technologies are used.
- Generally these counting detectors must themselves be very low background, which requires them to be clean and underground, and fabricated and shielded with low background materials.

### **Uranium Decay Chain**



### **Thorium Decay Chain**

Thorium A = 4n Gamma Intensities					13.52 1.600 16.2 0.72 12.75 0.304 15.5 0.16	Ra 228 5.75 a	63.823 0.264 204.68 0.021	Th 232 1.405x10 <sup>10</sup> a				
								911.204 25.8 968.971 15.8 338.320 11.27 964.766 4.99 463.004 4.40 794.947 4.25 209.253 3.89	Ac 228 6.15 h			
	238.632 43.3 300.087 3.28 115.183 0.592	Pb 212 10.64(1) h	804.9 0.0019	Po 216 145(2) ms	◀ 549.76 0.114	Rn 220 55.6(1) s	<b>∢</b> 240.986 4.10	Ra 224 3.66(4) d	84.373 1.220 215.983 0.254 131.613 0.131 166.410 0.104	Th 228 1.9116(16) a		
2614.533 99.0 583.191 84.5 510.77 22.6 860.564 12.42 277.351 6.31 763.13 1.81	Tl 208 3.053(4) m	α 39.858 1.091	• Bi 212 60.55(6) m 35.94% 64.06%	β 727.330 6.58 1620.50 1.49 785.37 1.102								
		Pb 208 stable		Po 212 299(2) ns								

### **Other Interesting Isotopes**



### Occasionally Present:

<sup>138</sup>La and <sup>176</sup>Lu

- <sup>54</sup>Mn at 834.85 keV Observed in Stainless Steel
- <sup>7</sup>Be at 477.60 keV Observed in Carbon based materials, due to neutron activation, samples are particularly affected after long flights.

Observed in rare earth samples such as Nd or Gd.